

Problem 1:

(1)

- Newtonian limit: we analyse static configuration \rightarrow time independent

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu} \quad \text{with } |h_{\mu\nu}| \ll 1$$

additionally $\vec{v} \ll 1 \rightarrow \frac{d\vec{x}}{d\tau} = \frac{dt}{d\tau} \underbrace{\frac{d\vec{x}}{dt}}_{=\vec{v}} \ll \frac{dt}{d\tau}$

- geodesic equation:

$$\frac{d^2 x^\mu}{d\tau^2} + \Gamma_{\alpha\beta}^\mu \frac{dx^\alpha}{d\tau} \frac{dx^\beta}{d\tau} = 0$$

In Newtonian limit:

$$\frac{d^2 x^i}{d\tau^2} + \Gamma_{00}^i \left(\frac{dt}{d\tau}\right)^2 = 0$$

$$\frac{d^2 x^i}{dt^2} = -\Gamma_{00}^i$$

2x chain rule

- Christoffel symbols:

$$\Gamma_{\alpha\beta}^\mu = \frac{1}{2} g^{\mu\nu} (\partial_\beta g_{\nu\alpha} + \partial_\alpha g_{\nu\beta} - \partial_\nu g_{\alpha\beta})$$

$$\Gamma_{00}^i \approx \frac{1}{2} \eta^{i\nu} (0 + 0 - \partial_\nu h_{00}) = -\frac{1}{2} \eta^{i\nu} \partial_\nu h_{00}$$

$$\Rightarrow \frac{d^2 x^i}{dt^2} = -\frac{1}{2} \partial_i h_{00} \quad (1)$$

$$\bullet R^S_{\sigma\mu\nu} = \partial_\mu \Gamma^S_{\nu\sigma} - \partial_\nu \Gamma^S_{\mu\sigma} + \underbrace{\Gamma^S_{\mu\lambda} \Gamma^\lambda_{\nu\sigma}}_{\sim \mathcal{O}(h^2)} - \underbrace{\Gamma^S_{\nu\lambda} \Gamma^\lambda_{\mu\sigma}}_{\sim \mathcal{O}(h^2)}$$

$$R^i_{\sigma j \sigma} \approx \partial_j \Gamma^i_{\sigma\sigma} + 0$$

$$R_{00} = R^i_{\sigma i \sigma} \approx -\partial_i \left(\frac{1}{2} \eta^{i\nu} \partial_\nu h_{00} \right) = \frac{1}{2} \nabla^2 h_{00} \quad (2)$$

• Einstein equation: $\sim \mathcal{O}(h)$

$$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} R \eta_{\mu\nu} = 8\pi G T_{\mu\nu}$$

$$R^\mu_{\mu} - \frac{1}{2} R \eta^\mu_{\mu} = 8\pi G T^\mu_{\mu}$$

$$R = -8\pi G T$$

$$\rightarrow R_{\mu\nu} = 8\pi G \left(T_{\mu\nu} - \frac{1}{2} T \eta_{\mu\nu} \right)$$

• With (1) and (2) we obtain for $\mu\nu=00$:

$$\nabla^2 V = 8\pi G \left(T_{00} - \frac{1}{2} T \right)$$

(2)

• Hilbert energy-momentum tensor:

$$T_{\mu\nu} = 2 \frac{\partial \mathcal{L}}{\partial g^{\mu\nu}} - g_{\mu\nu} \mathcal{L}$$

$$\begin{aligned} \bullet \mathcal{L} &= \frac{1}{2} (\partial_\mu \Phi)^2 - \frac{\lambda}{4} (\Phi^2 - v^2)^2 \\ &= \frac{1}{2} g^{\mu\nu} \partial_\mu \Phi \partial_\nu \Phi - \frac{\lambda}{4} (\Phi^2 - v^2)^2 \end{aligned}$$

$$= \frac{1}{2} g^{\mu\nu} \partial_\mu \Phi \partial_\nu \Phi - \frac{\lambda}{4} (\Phi^2 - v^2)^2$$

$$\rightarrow T_{\mu\nu} = \partial_\mu \Phi \partial_\nu \Phi - \frac{1}{2} g_{\mu\nu} \partial_\alpha \Phi \partial^\alpha \Phi + g_{\mu\nu} \frac{\lambda}{4} (\Phi^2 - v^2)^2$$

$$\bullet \Phi_{DW} = v \cdot \tanh\left(\frac{1}{2} m_h x\right) \quad \text{with} \quad m_h = 2\sqrt{2\lambda} v$$

$$T_{00} = \frac{1}{2} (\partial_x \Phi_{DW})^2 + \lambda (\Phi_{DW}^2 - v^2)^2 = \epsilon_{DW}$$

$$T_{11} = \frac{1}{2} (\partial_x \Phi_{DW})^2 - \lambda (\Phi_{DW}^2 - v^2)^2 = 0$$

by e.o.m.

$$T_{22} = -\frac{1}{2} (\partial_x \Phi_{DW})^2 - \lambda (\Phi_{DW}^2 - v^2)^2 = -\epsilon_{DW} = T_{33}$$

$$\rightarrow T_{\mu\nu} = \epsilon_{DW} \text{diag}(1, 0, -1, -1)$$

(3)

$$\bullet T = g^{\mu\nu} T_{\mu\nu} = T_{00} - T_{ii}$$

$$= 3\epsilon_{DW}$$

$$\bullet \vec{\nabla}^2 V = 8\pi G \left(\epsilon_{DW} - \frac{3}{2} \epsilon_{DW} \right)$$

$$= -4\pi G \epsilon_{DW}$$

→ Because of the minus sign the gravitational field of the domain wall is repulsive

(4)

$$\bullet \epsilon_{DW} = \frac{1}{2} (\partial_x \Phi_{DW})^2 + \lambda (\Phi_{DW}^2 - v^2)^2$$

e.o.m.?

$$= 2\lambda (\Phi_{DW}^2 - v^2)^2$$

$$\bullet V(\vec{r}) = \int \left(-\frac{1}{4\pi|\vec{r}-\vec{r}'|} \right) \cdot (-8\pi G\lambda (\phi_{0w}^2 - v^2)^2) d^3r'$$

without loss of
generality \rightarrow
we can write \rightarrow
 $V(x,y,z) = V(x,0,0)$

$$= 2G\lambda v^4 \cdot 2\pi \int dx' \int_0^\infty ds' \frac{s'}{\sqrt{s'^2 + (x-x')^2}} \left(\tanh\left(\frac{1}{2} m_h x'\right)^2 - 1 \right)^2$$

shift the potential

$$= \frac{\pi}{2} G m_h^2 v^2 \int dx' \left(\tanh\left(\frac{1}{2} m_h x'\right)^2 - 1 \right)^2 (\infty + |x' - x|)$$

$$= \frac{\pi}{2} G v^2 \int d(m_h x') \left(\tanh\left(\frac{1}{2} m_h x'\right)^2 - 1 \right)^2 |m_h x' - m_h x|$$

$$= \frac{\pi}{2} G v^2 \left[\int_{-\infty}^{m_h x} d(m_h x') \left(\tanh\left(\frac{1}{2} m_h x'\right)^2 - 1 \right)^2 (m_h x - m_h x') \right. \\ \left. + \int_{m_h x}^{\infty} d(m_h x') \left(\tanh\left(\frac{1}{2} m_h x'\right)^2 - 1 \right)^2 (m_h x' - m_h x) \right]$$

Mathematical

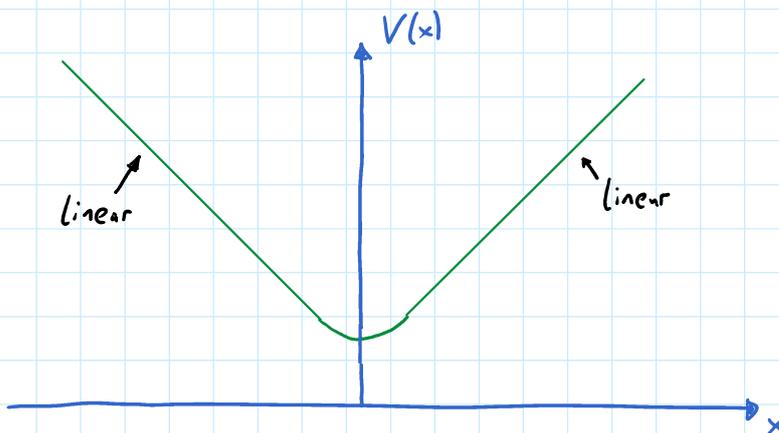
$$= \frac{2\pi}{3} G v^2 \left(\ln(16) + 4 \ln\left(\cosh\left(\frac{m_h x}{2}\right)\right) - \frac{2}{1 + \cosh(m_h x)} \right)$$

$$\bullet \ln\left(\cosh\left(\frac{x}{2}\right)\right) = \ln\left(\frac{1}{2} (e^{-m_h x/2} + e^{m_h x/2})\right)$$

$$\xrightarrow{|x| \rightarrow \infty} \ln\left(\frac{1}{2}\right) + \frac{m_h |x|}{2}$$

$$\frac{2}{1 + \cosh(m_h x)} \xrightarrow{|x| \rightarrow \infty} 0$$

$$\Rightarrow V(x) \xrightarrow{|x| \rightarrow \infty} \frac{\pi}{3} G v^2 m_h |x| + \frac{2\pi}{3} G v^2 \ln 8$$



Problem 2:

$$\bullet \mathcal{L} = \frac{1}{2} \partial_r \Phi^a \partial^r \Phi^a - V(\Phi) \quad , \quad d+1 \text{ dimensions}$$

$a=1, \dots, n$
↓

energy $E = T + U$ with $T = \int d^d x \frac{1}{2} \partial_i \Phi^a \partial_i \Phi^a$
 $U = \int d^d x V(\Phi)$

is minimized by a soliton solution.

(1)

$$\bullet \Phi_\lambda^a(x) = \Phi^a(\lambda x)$$

$$T \mapsto \int d^d x \frac{1}{2} \frac{\partial}{\partial x^i} \Phi^a(\lambda x) \frac{\partial}{\partial x^i} \Phi^a(\lambda x)$$

$$= \lambda^{2-d} \int d^d(\lambda x) \frac{1}{2} \frac{\partial}{\partial(\lambda x^i)} \Phi^a(\lambda x) \frac{\partial}{\partial(\lambda x^i)} \Phi^a(\lambda x)$$

$$= \lambda^{2-d} T$$

$$U \mapsto \int d^d x V(\Phi(\lambda x)) = \lambda^{-d} \int d^d(\lambda x) V(\Phi(\lambda x)) = \lambda^{-d} U$$

$$\rightarrow E(\lambda) = \lambda^{2-d} T + \lambda^{-d} U$$

(2)

• The soliton extremises the energy
and $\Phi_{cl}^a(\lambda x)$ is a soliton solution for $\lambda=1$.

(3)

$$\bullet \frac{dE}{d\lambda} = (2-d) \lambda^{1-d} T - d \lambda^{-d-1} U$$

$$\left. \frac{dE}{d\lambda} \right|_{\lambda=1} \stackrel{!}{=} 0$$

$$\rightarrow (2-d) T - d U = 0$$

- First, notice that $T, U \geq 0$

For $d > 2$: $-12-dT = dU$

→ only possible solution is $T = U = 0$

→ trivial vacuum

- for $d = 2$: $U = 0$

→ trivial vacuum

- for $d = 1$: $T = U$

→ allowed with soliton solution

(4)

- $\mathcal{L} = -\frac{1}{2} \text{Tr} (G_{\mu\nu} G^{\mu\nu}) + \text{Tr} ((D_\mu \Phi)^\dagger (D^\mu \Phi)) - V(\Phi)$

scale transformation: $\Phi_\lambda(x) = \Phi_c(\lambda x)$

$$A_\lambda^i(x) = \lambda A_c^i(\lambda x)$$

- potential scales as before:

$$U \mapsto \lambda^{-d} U$$

- $T = \int d^d x \text{Tr} ((D_i \Phi(x))^\dagger D_i \Phi(x))$

$$\mapsto \int d^d x \text{Tr} ((\partial_i \Phi^\dagger(\lambda x) - g \lambda [A_i^\dagger(\lambda x), \Phi^\dagger(\lambda x)]) \cdot (\partial_i \Phi(\lambda x) + g \lambda [A_i(\lambda x), \Phi(\lambda x)]))$$

$$= \lambda^{2-d} \int d^d(\lambda x) \text{Tr} \left(\left(\frac{\partial}{\partial(\lambda x^i)} \Phi^\dagger(\lambda x) - g [A_i^\dagger(\lambda x), \Phi^\dagger(\lambda x)] \right) \cdot \left(\frac{\partial}{\partial(\lambda x^i)} \Phi(\lambda x) + g [A_i(\lambda x), \Phi(\lambda x)] \right) \right)$$

$$= \lambda^{2-d} T$$

$$\bullet G_{ij} = \partial_i A_j - \partial_j A_i + [A_i, A_j]$$

$$\mapsto \partial_i \lambda A_j(\lambda x) - \partial_j \lambda A_i(\lambda x) + \lambda^2 [A_i(\lambda x), A_j(\lambda x)]$$

$$= \lambda^2 \left[\frac{\partial}{\partial(\lambda x^i)} A_j(\lambda x) - \frac{\partial}{\partial(\lambda x^j)} A_i(\lambda x) + [A_i(\lambda x), A_j(\lambda x)] \right]$$

$$\rightarrow G = \int d^d x \frac{1}{2} \text{Tr}(G_{ij} G_{ij})$$

$$\mapsto \lambda^{4-d} G$$

$$\bullet E(\lambda) = \lambda^{4-d} G + \lambda^{2-d} T + \lambda^{-d} U$$

$$(5) \bullet \left. \frac{dE}{d\lambda} \right|_{\lambda=1} = (4-d)G + (2-d)T - dU \stackrel{!}{=} 0$$

Note that $G, T, U \geq 0$

$$\bullet d \geq 4: \quad +|4-d|G + |2-d|T + dU = 0$$

only $G=T=U=0$ is allowed

$$\bullet d < 4: \quad |4-d|G + |2-d|T = dU$$

Also non-trivial solutions are allowed

For $d=1, 2, 3$ we can have solitons.

Problem 3:

$$(1) \bullet \langle |\Phi| \rangle = v \quad \rightarrow \quad \Phi = (\varphi + v) e^{i\theta}$$

$$\bullet \mathcal{L} \supset \lambda ((\varphi + v)^2 - v^2)^2$$

- $$\mathcal{L} \supset \lambda (\langle \Phi \rangle^2 - v^2)^2$$

$$= \lambda (\varphi^2 + 2v\varphi)^2 \supset 4\lambda v^2 \varphi^2 = \frac{1}{2} (8\lambda v^2) \varphi^2$$

$$\rightarrow m_h = 2\sqrt{2\lambda} v$$

- $$D_\mu \langle \Phi \rangle = ig A_\mu v$$

$$\rightarrow \mathcal{L} \supset (D_\mu \langle \Phi \rangle)^* (D^\mu \langle \Phi \rangle) = g^2 v^2 A_\mu A^\mu$$

$$\rightarrow m_v = \sqrt{2} g v$$

(2)

- $$E = H$$

for static case

$$= - \int d^2x \mathcal{L}$$

$$= \int d^2x \left(\frac{1}{4} F_{ij} F_{ij} + (D_i \Phi)^* (D^i \Phi) + \lambda (\langle \Phi \rangle^2 - v^2)^2 \right)$$

- conditions for finite energy:

$$D_i \Phi \xrightarrow{r \rightarrow \infty} 0$$

$$|\Phi| \xrightarrow{r \rightarrow \infty} v$$

(3)

- $$\Phi \xrightarrow{r \rightarrow \infty} v \cdot e^{in\theta}$$

$$D_i \Phi \xrightarrow{r \rightarrow \infty} v e^{in\theta} \text{ in } \partial_i \theta + igv A_i e^{in\theta} \stackrel{!}{=} 0$$

$$\rightarrow A_i \xrightarrow{r \rightarrow \infty} \frac{1}{g} n \partial_i \theta$$

- $$\theta = \arctan \frac{y}{x}$$

$$\partial_x \theta = \frac{1}{1 + \frac{y^2}{x^2}} \cdot \left(-\frac{y}{x^2} \right) = -\frac{y}{x^2 + y^2} = -\frac{y}{r^2}$$

$$\partial_x \Theta = \frac{-y}{1+y^2} \cdot \left(-\frac{y}{x^2}\right) = -\frac{y}{x^2+y^2} = -\frac{y}{r^2}$$

$$\partial_y \Theta = \frac{1}{1+y^2} \cdot \frac{1}{x} = \frac{x}{x^2+y^2} = \frac{x}{r^2}$$

$$\rightarrow A_i \xrightarrow{r \rightarrow \infty} -\frac{1}{g} n \epsilon_{ij} \frac{r^j}{r^2}$$

(4)

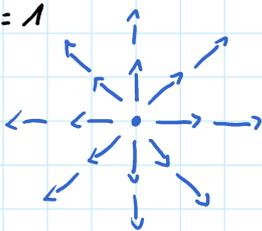
• magnetic flux:

$$\begin{aligned} \Phi &= \int_S d\vec{S} \cdot \vec{B} = \oint_{\partial S} d\vec{l} \cdot \vec{A} \\ &= \frac{n}{g} \oint_{\partial S} (\partial_i \Theta) dx^i \\ &= \frac{n}{g} \int_0^{2\pi} d\Theta = \frac{2\pi n}{g} \end{aligned}$$

(5)

$$\bullet \begin{pmatrix} \operatorname{Re} \hat{\Phi} \\ \operatorname{Im} \hat{\Phi} \end{pmatrix} = \begin{pmatrix} \cos(n\Theta) \\ \sin(n\Theta) \end{pmatrix}$$

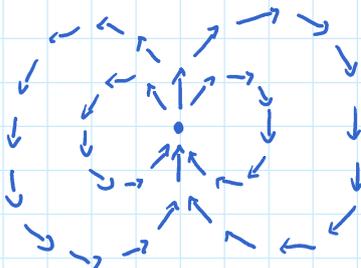
$n=1$



$n=-1$

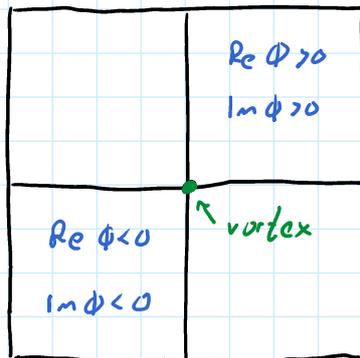


$n=2$

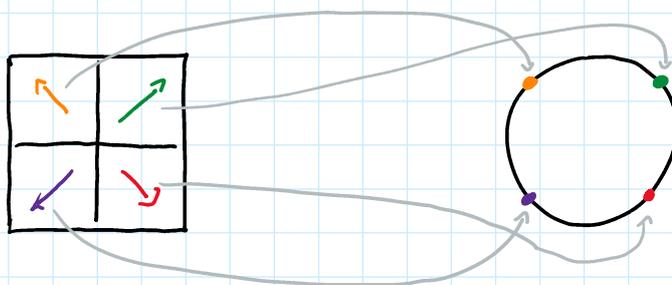


(6)

- We can have many causally disconnected areas that obtain a different vacuum expectation values. Note that in the given theory we have infinitely many vacuum expectation values.
- Imagine that one area obtains a value with $\text{Re } \Phi, \text{Im } \Phi > 0$ and another area obtains a value with $\text{Re } \Phi, \text{Im } \Phi < 0$, then there is always a point where the field Φ has to vanish. There, a vortex will appear:



- Around the vortex each point can be mapped to a point on the vacuum manifold:



└

Comment: A stable domain wall cannot exist in

Comment: A stable domain wall cannot exist in this theory. We will explain the reason on a later sheet.

(7)

- Yes, there can be a string solution.

"Instead of one vortex we have many vortices layered on top of each other."

