

GUT Course 22/23

Lecture X

25/11/2022

LHV

2022



# Monopoles in Cosmology

$$\partial_\mu F^{\mu\nu} = j^\nu \Rightarrow \partial_\mu j^\mu = 0$$

$$\partial_\mu \tilde{F}^{\mu\nu} = \textcircled{j^\nu} \Rightarrow \partial_\mu j^\mu = 0$$



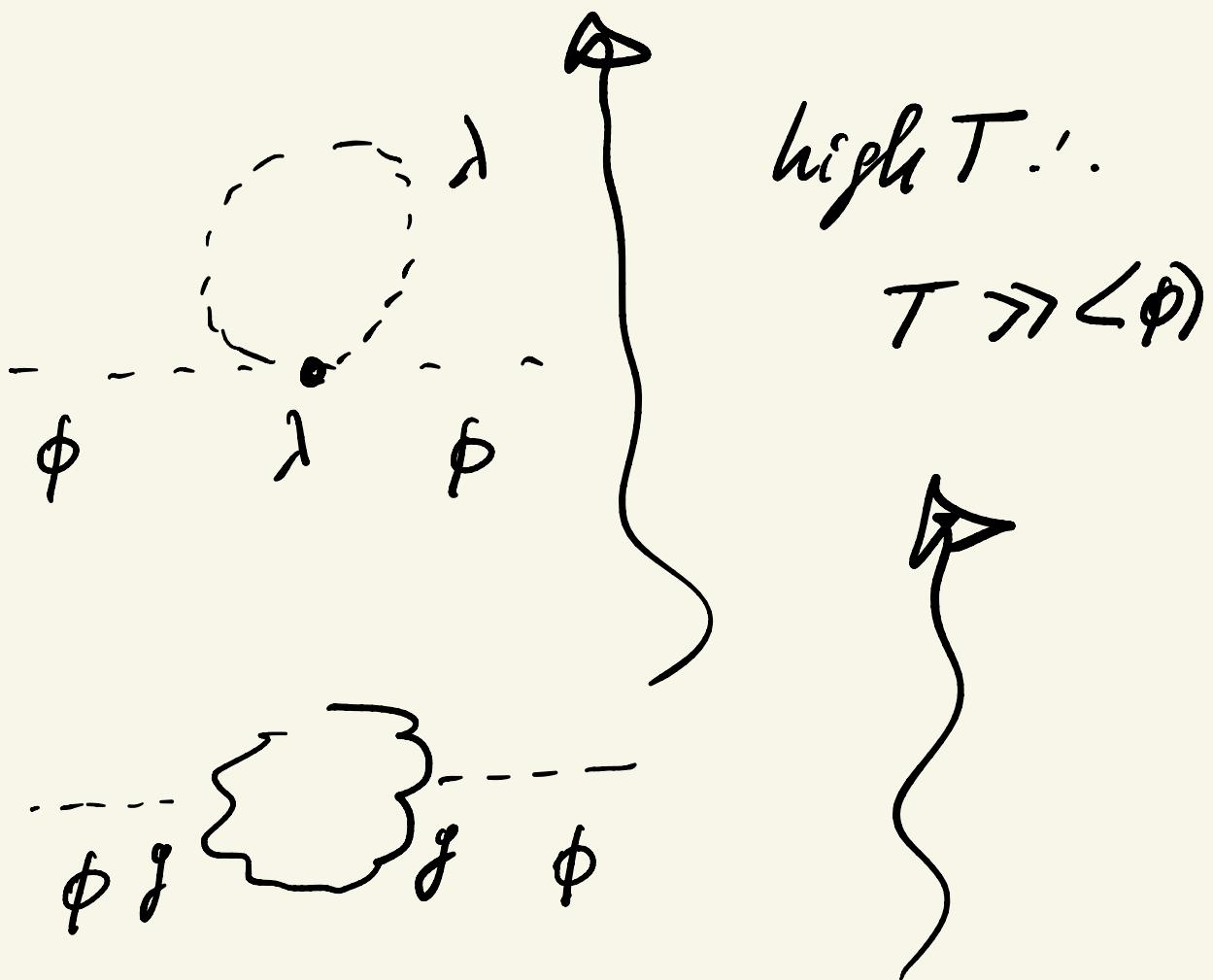
magnetic current

- at high  $T \Rightarrow \langle \phi \rangle = 0$



$$V(T) = V(0) + aT^2 \phi^\top \phi$$

$$a = \lambda + g^2 > 0 \quad (\lambda > 0)$$



$$\Rightarrow \mu_\phi^2 = \cancel{\mu_0^2} + \alpha T^2 > 0$$

correlation length  $\approx$  horizon



$\sim 1$  monopole/horizon

$$\bullet \boxed{d_H = t = \frac{H_{\text{pc}}}{T^2}}$$

a)  $T \rightarrow 0, t \rightarrow \infty$

b)  $H_{\text{pc}} \rightarrow \infty$  ( $\alpha_N \rightarrow 0$ )  $\Rightarrow$   
 $t \rightarrow \infty$

$$H \equiv \dot{R}/R \Rightarrow H^2 = \alpha_N \rho$$

$T \gg \langle \phi \rangle \Rightarrow \rho \sim T^4$

$$\Rightarrow H = \frac{T^2}{H_{\text{pc}}}$$

$H = 1/t \Rightarrow \boxed{t = \frac{H_{\text{pc}}}{T^2}}$

early universe (confirmed)

$$T \sim 10^{-100} \text{ eV}$$

Black Holes :

$$R_{BH} = G_N M_{BH}$$

Hawking :

$$G_N \rightarrow 0, M_{BH} \rightarrow \infty$$

$$\therefore R_{BH} = \text{finite}$$



$$\rho_M = \left(\frac{1}{d_H}\right)^3$$

Kibble  
density of monopoles medium

$$u_H = \frac{T^6}{M_P^3}$$

$$T \sim \langle \phi \rangle$$

$$u_\gamma = T^3$$

$$u_B - u_{\bar{B}} = 10^{-10} u_\gamma$$

## UNIVERSE

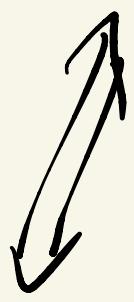
$$N_\gamma = \text{const}$$

$$N_B - N_{\bar{B}} = \text{const}$$

$$\Rightarrow N_\gamma = u_\gamma R^3$$

$$\Rightarrow RT = \text{const}$$

$$R = R_0$$



big-bang

total # is conserved



$$\frac{M_H}{M_B} \simeq 10^{10} \frac{T^3}{M_P^3}$$

$\vartheta = \langle \phi \rangle$

$M_H \sim \frac{1}{g} \vartheta$

$$\Rightarrow \frac{M_H}{M_B} = \vartheta / \text{GeV} \quad 10^{10} \frac{T^3}{M_P^3}$$

$\eta$

universe  $= \vartheta^4 (\text{GeV}) \quad 10^{10} / M_P^3 \text{ GeV}^3$

*big-bang*

$$M_p \simeq 10^{19} \text{ GeV}$$

$$\frac{M_H}{M_B} \leq O(1)$$



$$10^{10} \vartheta^4(\text{GeV}) \leq 10^{5+}$$

$$\vartheta \leq 10^{12} \text{ GeV}$$

$$\text{GUT} \Leftrightarrow \vartheta_{\text{GUT}} \simeq 10^{16} \text{ GeV}$$

$$\Rightarrow \frac{M_H^{\text{GUT}}}{M_B} \simeq 10^{16}$$

monopole  
problem  
in GUT

However, if  $\vartheta \lesssim 10^{12} \text{ GeV}^-$

$\Rightarrow DM ?$

Pati-Salam

$g-l$  unification

$\vartheta_{PS}$  could be  $\leq 10^{12} \text{ GeV}$

$$G_{PS} = SU(2)_L \times SU(2)_R \times SU(4)_C$$



monopoles

$GUT =$  Grand Unification



Single gauge non-Abelian

group



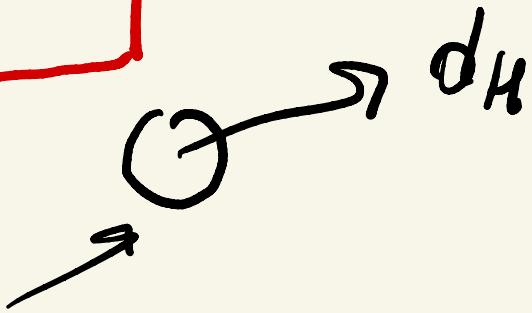
$G_{SM} \subseteq SU(5), SO(10)$

$E(6)$

Ways out of

monopole problem:

• inflation



$T^{\circ}, t^{\circ}$  : from a horizon

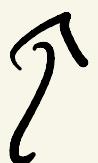


# of monopoles  $\sim \Theta(1)$

in our universe



big - bang



$$d_H = \frac{M_P}{T^2} \quad R = \frac{c}{T}$$

today :  $T_0 \simeq 10^{-13} \text{ GeV}$

$$R_0 \simeq 10^{29} \text{ cm}$$

$$\text{cm GeV} \simeq 10^{14}$$

$$\Rightarrow \boxed{C_0 = C = 10^{30}}$$

$$d_H' = R_0$$



$$\left(\frac{R}{d_H}\right)^3 = \left(10^{30} \frac{T}{M_P}\right)^3$$

$$T \simeq \text{MeV} \Rightarrow \left(\frac{R}{d_H}\right)^3 \simeq 10^{24}$$

nucleosynthesis

## HORIZON PROBLEM

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S & B at high T

$\phi_1, \phi_2 \leftarrow$  real scalar

$$V(0) = -\frac{\mu_i^2}{2} \phi_i^2 + \frac{\lambda_i}{4} \phi_i^4 + \frac{\lambda_3}{2} \phi_1^2 \phi_2^2$$

Symmetry (ies) ?  $(\mu_i^2 > 0)$

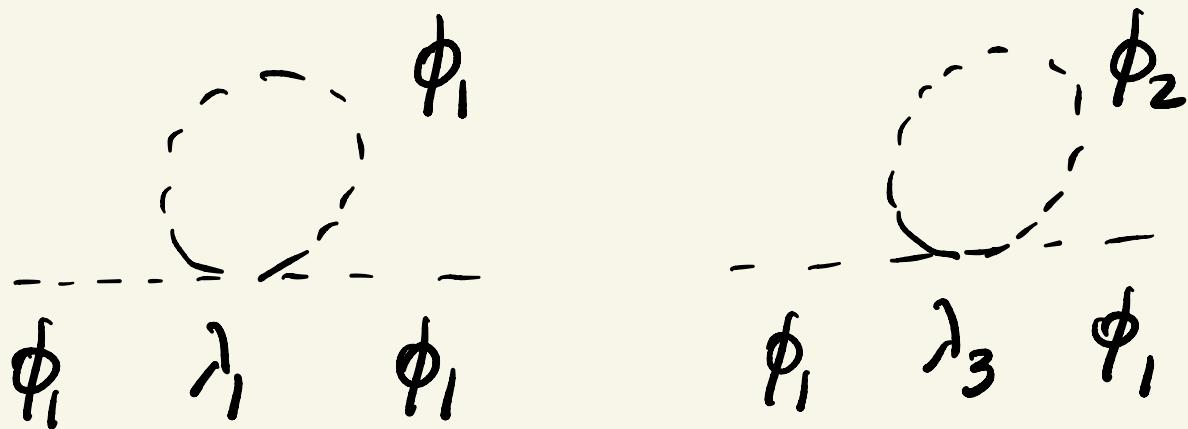
$$D_1 : \phi_1 \rightarrow -\phi_1, \phi_2 \rightarrow \phi_2$$

$$D_2: \quad \phi_1 \rightarrow \phi_1, \quad \phi_2 \rightarrow -\phi_2$$


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$$T=0: \quad \langle \phi_1 \rangle \neq 0 \neq \langle \phi_2 \rangle$$

high  $T \therefore T \gg \langle \phi_i \rangle$



+ 1  $\leftrightarrow$  2 diagrams

$$V(T) = V(0) + \alpha_i T^2 \phi_i^2$$

$$\alpha_1 = \lambda_1 + \lambda_3, \quad \alpha_2 = \lambda_2 + \lambda_3$$

$V(0)$  = bounded



$$\lambda_1, (\lambda_1, \lambda_2) > 0$$

$$\lambda_1, \lambda_2 > \lambda_3^2 \leftarrow \text{DERIVE}$$



$\lambda_3$  can be negative

$$a_1 = \lambda_1 + \lambda_3 < 0$$

possible

$$\Rightarrow a_2 = \lambda_2 + \lambda_3 > 0$$



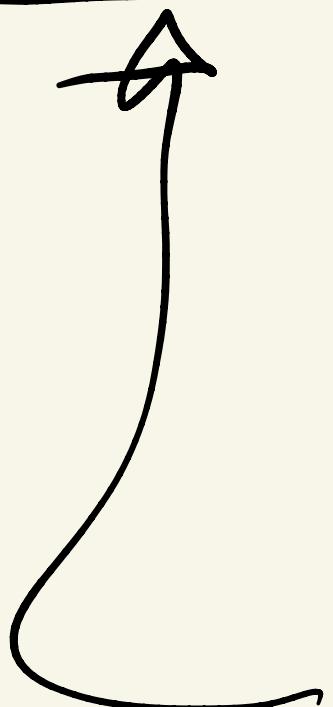
$$-\left| \alpha_1 \right| T^2 \phi_1^2 \Rightarrow$$

$$\langle \phi \rangle_T \neq 0$$



$$\langle \phi \rangle_T \simeq T$$

Holopatwa, 6.8.  
78



Weinberg '74  
(told by Coleman)

# Rochelle salt

$T_{1c} \rightarrow$  more crystallized  
 $T_{2c} (\gg T_{1c}) \Rightarrow$  melts

QFT systems  $\sim$

Dvali, Melville, S.  
'95



rid of monopole problem

Dveli, G.S. '94

grid of domestic wells