

Neutrino Physics Course

Lecture XXIII

14/7/2023

LMO

Summer 2023



$d = 5$ $LN\bar{V}$ interaction (II)

1979 Weinberg

neutrino mass

$$\frac{1}{\Lambda_{\text{new}}} \bar{\nu} \nu \phi_0 \phi_0 \longrightarrow \frac{1}{\Lambda_{\text{new}}} \bar{l} l \phi \phi$$

$$l = \begin{pmatrix} \nu \\ e \end{pmatrix}_L$$

$$\phi = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix}$$

$\Delta L \neq 0$ motivation

$l l \leftarrow$ at least two

↑

$$l^T i \sigma_2 c l = 0$$

$l \ l \ \phi \ \phi$

$$\left. \begin{array}{l} \Delta y = 0 \\ \Delta T_3 = 0 \end{array} \right\}$$

$$\underbrace{\hspace{10em}}_S \quad \underbrace{\hspace{10em}}_{S = \text{Huglet}}$$
$$(l^T i \sigma_2 \phi) \ C \ (\phi^T i \sigma_2 l)$$

(I)

Λ_{new}

$$(II) \quad (l^T i \sigma_2 \bar{\sigma} \phi) \ C \ (\phi^T i \sigma_2 \bar{\sigma} l) / \Lambda_{new}$$

$$\overbrace{\quad\quad\quad}^{\vec{V}(T)}$$

$$\overbrace{\quad\quad\quad}^{\vec{V}(T)}$$

// //

vector

triplet

$$S = l^T i \sigma_2 \phi = \text{fermion}$$

$$\vec{V} = l^T i \sigma_2 \vec{\sigma} \phi = \text{---}$$

$$(II) \quad \underline{(l^T C i \sigma_2 \vec{\sigma} l) (\phi^T i \sigma_2 \vec{\sigma} \phi)}$$



$$l^T \dots i \sigma_2 \vec{\sigma} l = \text{scalar}$$

$$\phi^T \sigma_1 \sigma_2 \bar{\sigma} \phi = -11 -$$

$$\rightarrow \phi_{\text{new}} = \begin{pmatrix} 0 \\ \phi_0 \end{pmatrix} \Rightarrow$$

$$\text{(II)} \rightarrow \# \frac{v^T c v \phi_0 \phi_0}{\Lambda_{\text{new}}}$$

\Downarrow

I, II, III = equal

(proportional to each other)

\Uparrow

1
 in the general $\phi = \begin{pmatrix} \phi^+ \\ \phi_0 \end{pmatrix}$

$$\underline{I} \sim \underline{II} \sim \underline{III} \propto$$

$$\left(\begin{array}{cc} \nu_L^T c \nu_L \phi_0 \phi_0, & \nu_L^T c e_L \phi_0 \phi^+, \\ e_L^T c e_L \phi^+ \phi^+ \end{array} \right) \frac{1}{\Lambda_{new}}$$

↓ dust settles

$$\nu_L \nu_L \frac{M_{SM}^2}{\Lambda_{new}} \Leftrightarrow \nu = \text{Majorana}$$

$$L_{\text{int}} = \frac{\nu \nu \phi_0 \phi_0}{\Lambda_{\text{new}}}$$

$$\phi_0 = h + \cancel{t\phi}$$



$$\frac{\nu \nu \nu_{SM} h}{\Lambda_{\text{new}}}$$

$$\frac{\nu \nu \nu_{SM}^2}{\Lambda_{\text{new}}}$$



$$\bar{f} f \frac{m_f}{\nu_{SM}} h$$

$$\boxed{m_\nu = \frac{\nu_{SM}^2}{\Lambda_{\text{new}}}} \quad (*)$$

(f = charged fermion)

$$\Leftrightarrow Y_f = \frac{m_f}{\nu_{SM}}$$

⇓

$$\Gamma(h \rightarrow f\bar{f}) \propto \frac{w_f^2}{v_{SM}^2} M_h$$

⇕ analogy with J

$$\nu \nu \frac{v_{SM}}{\Lambda_{new}} h$$

$$\Rightarrow g_\nu = \frac{v_{SM}}{\Lambda_{new}}$$

but: $\left| \frac{w_\nu}{v_{SM}} = \frac{v_{SM}}{\Lambda_{new}} \right. \quad (*)$

$$g_{\nu} = \frac{m_{\nu}}{v_{SM}}$$

$$\Rightarrow \Gamma(h \rightarrow \nu\nu) \propto \frac{m_{\nu}^2}{v_{SM}^2} m_h$$

$$m_{\nu} < 1 \text{ eV}$$

$$v_{SM} \simeq 100 \text{ GeV} \simeq 10^5 \text{ eV}$$

$$\Rightarrow B_{\nu}(h \rightarrow \nu\nu) \equiv \frac{\Gamma(h \rightarrow \nu\nu)}{\Gamma_{\text{total}}}$$

$$B_{\nu}(h \rightarrow \nu\nu) \leq 10^{-22}$$

$d=5$ implies (by itself)

no new physics

(except for $\Delta L=2$)

(I) \Leftrightarrow Type I seesaw

(II) \Leftrightarrow -||- II -||-

(III) \Leftrightarrow -||- III -||-

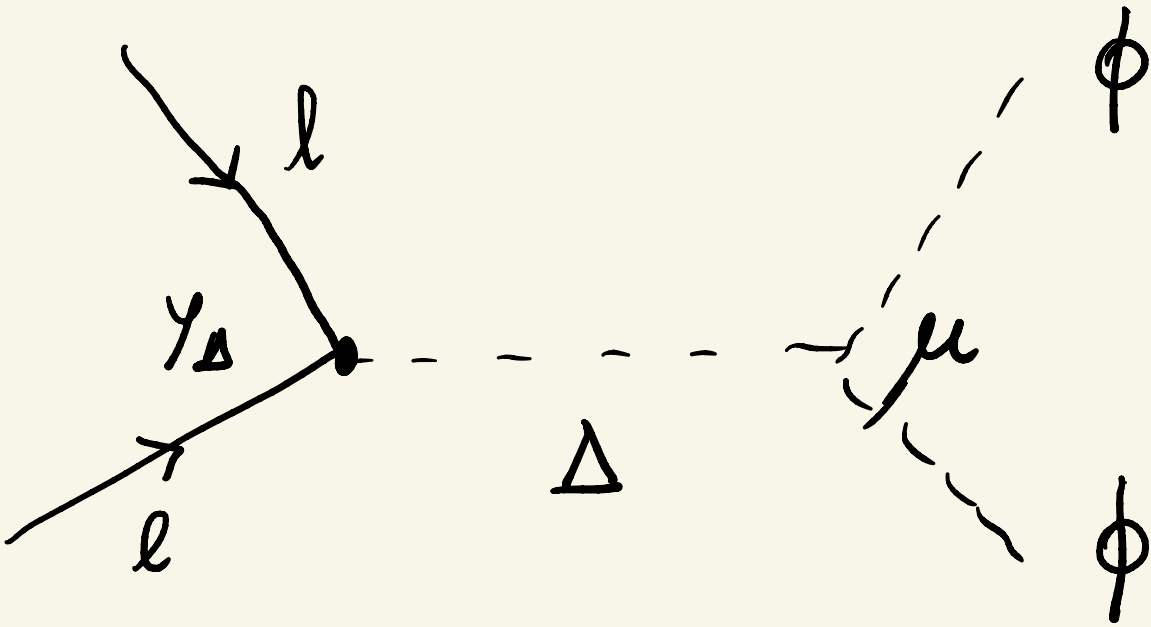
Example

Type II seesaw

$SU(2)_L$ triplet

$$Y_\Delta \ell \Delta \ell + \mu \phi \Delta^* \phi$$

$$Y(\Delta) = 2$$



$$\mathcal{L}_{\text{eff}} = \gamma_{\Delta\mu} \ell \ell \frac{1}{k^2 - m_{\Delta}^2} \phi \phi$$

$$\xrightarrow{(k \rightarrow 0)} \gamma_{\Delta\mu} \ell \ell \frac{1}{m_{\Delta}^2} \phi \phi$$

$$\mathcal{L}_{\text{eff}}^{(d=5)} \text{ (Weyly)} = \ell \ell \phi \phi \frac{1}{\Lambda_{\text{new}}}$$

$$\text{iff } \frac{1}{\Lambda_{\text{new}}} = \frac{m_{\Delta}^2}{\gamma_{\Delta\mu}}$$

Type I error

$$\frac{1}{\Lambda_{new}} = \frac{\mu_{NS}}{\gamma_D^2}$$

(here $d=5$ layers)
is better)

$$\Lambda_{new} = ?$$

Information ???

\Downarrow $d=5$

$$m_\nu = \frac{v_{SM}^2}{\Lambda_{new}} \quad m_\nu < 10^{-9} \text{ GeV}$$

$$\Rightarrow \Lambda_{new} = \frac{v_{SM}^2}{m_\nu}$$

$$\Lambda_{new} \gtrsim \frac{10^4}{10^{-9}} \text{ GeV}$$

$$\Lambda_{new} \gtrsim 10^{13} \text{ GeV}$$

Cheat

$$L_{eff} = \frac{ll \phi \phi}{\Lambda_{new}} Y_{eff}$$



$$\frac{1}{\Lambda_{\text{new}}} = \frac{g_{\text{eff}}}{\Lambda_{\text{new}'}}$$

$$\Lambda_{\text{new}'} = \Lambda_{\text{new}} g_{\text{eff}}$$

Type I seesaw : $g_{\text{eff}} = g_0^2$

$$m_0 \simeq m_e \Leftrightarrow g_0 \simeq g_e$$

$$g_e \simeq 10^{-5} \dots$$

$$\Rightarrow \Lambda_{\text{new}'} = 10^{-10} \quad \Lambda_{\text{new}} = 10^3 \text{ GeV}$$

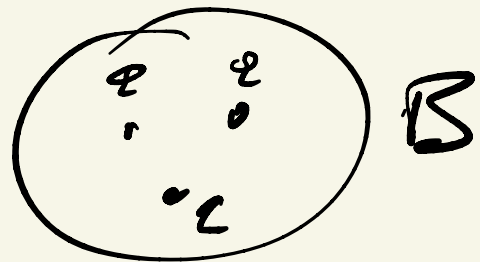
$d=5$ eff \rightarrow seesaw = UV
 completion

$\bar{U}\bar{V}$ = Ultra Violet

$\Delta L \neq 0$ $ll \phi\phi / \Delta$

$\Delta B \neq 0$

$$B_\phi = \frac{1}{3}$$



~~$q_k q_p$~~

~~$$3_c \times 3_c = 6_c + 3_c^*$$~~

~~quark conf.~~

~~(S) (A)~~

~~$q_e (B=0)$~~

$$\sum_{\alpha\beta\gamma} q_\alpha q_\beta q_\gamma = \text{color triplet}$$

\Downarrow

$$\Delta B \neq 0: \quad \underbrace{qqq l}_{s = \frac{1}{2} + \frac{1}{2} + \frac{1}{2} + \frac{1}{2} = 0}$$

Lorentz inv.

$SU(2) \times U(1)$ inv.

$$\Rightarrow \mathcal{L}_{\text{eff}} (\Delta B \neq 0) = \frac{qqq l}{\Lambda_{\text{new}}^2}$$

($\Lambda_{\text{new}} \gg M_W$)

p -neutrino decays

G.S. 2009

⇒ predictions



$$\tau_p \approx 10^{34} \text{ yr} \Rightarrow$$

$$\Lambda_{\text{new}} > 10^{15} \text{ GeV}$$

$$\tau_\nu \approx 10^{10} \text{ yr}$$

~ 1950 Maurice Goldhaber

"fed it in my
bones"

$$\tau_p \gtrsim 10^{18} \text{ yr}$$

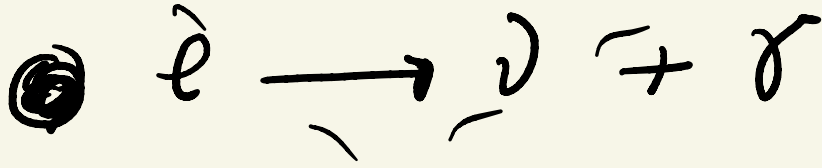


What is a "good" $\Delta L \neq 0$
operator?

• $d=5$ \rightarrow $ll \phi\phi$

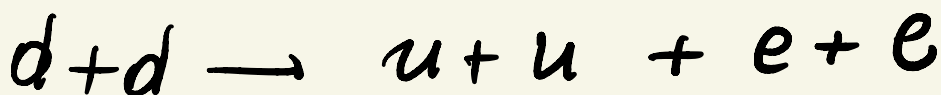
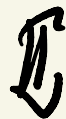
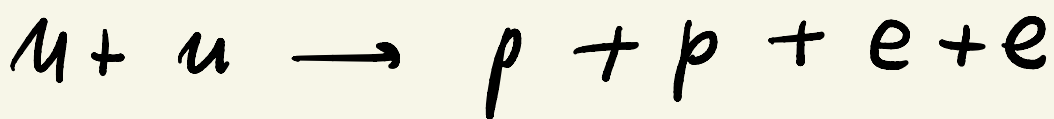
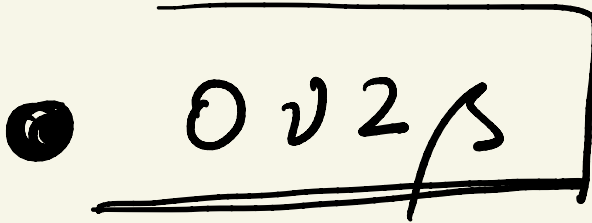
- "no good" : $\mu_\nu = M_{\text{GUT}}$

• $\Delta L = 2$ low energy process



$\Delta Q \neq 0$

$\tau_e \geq 10^{26} \text{ yr}$

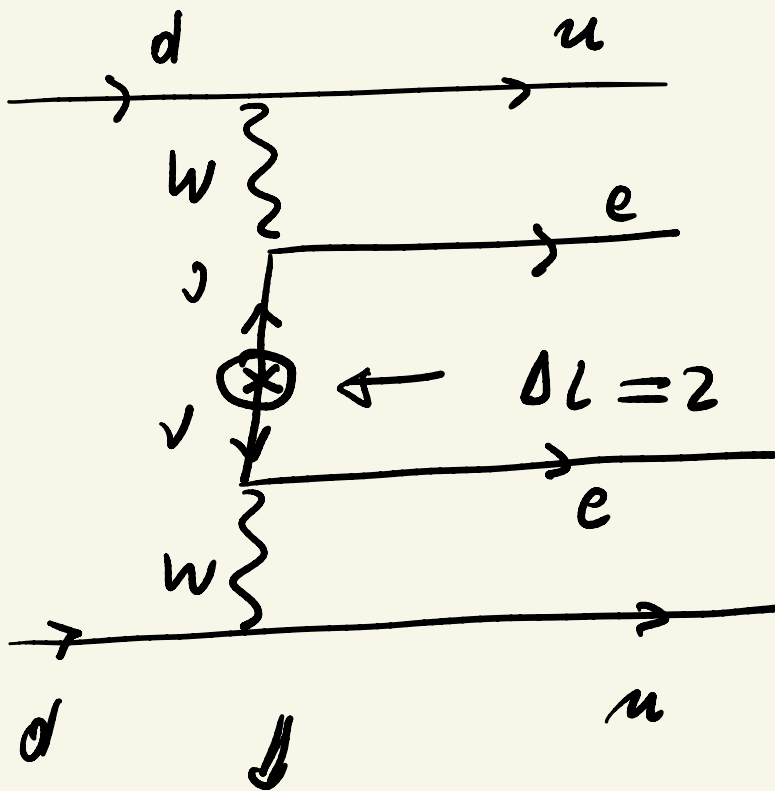


6 fermions ($\Delta_f = 3/2$)

$$L_{int} (d=g) = \frac{\bar{e} \bar{e} \bar{u} \bar{u} d d}{N_{new}^5}$$

$$\Delta L = 2$$

$$\Delta B = 0$$



$$G_F \approx 10^{-5} \text{ GeV}^{-2}$$

$$\bar{u} \bar{u} \bar{e} \bar{e} d d \quad G_F^2 \frac{M_\nu}{p^2}$$

$$(p \approx 100 \text{ GeV})$$

$$\frac{1}{\Lambda_{\text{new}}^5} \approx \frac{10^{-10} 10^{-10}}{10^{-2}} \text{ GeV}^{-5}$$

$$\approx 10^{-18} \text{ GeV}^{-5}$$

$$\Lambda_{\text{new}} \gtrsim 3 \times 10^3 \text{ GeV}$$

LHC??

$$\frac{\bar{u} \bar{u} d d \bar{e} \bar{e}}{\Lambda_{\text{new}}^5} \quad y_{\text{eff}} \quad \text{perturbativity}$$

$$(y_{\text{eff}} \leq O(1))$$

