

Neutrino Physics Course

Lecture XVIII

27/6/2023

LMU

Summer 2023

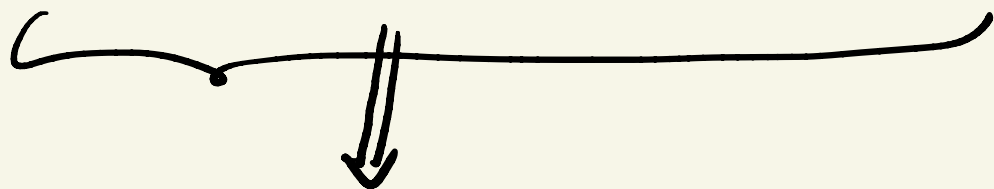


S.M Higgs in LRSM

$$\begin{array}{ccc} G_{LR} & \longrightarrow & G_{SM} & \longrightarrow & U_{em}(1) \\ \langle \Delta_R \rangle \equiv v_R & & \langle \Phi \rangle & & \\ \parallel & & \parallel & & \\ M_R & & M_W (M_L) & & \end{array}$$

$$M_R \gg M_L (M_W)$$

$$V = -V_\Delta + V_\Phi + V_{\Phi\Delta}$$



mass spectrum of Φ
(leading terms)

$$\begin{aligned}
V_{\Phi \Delta} &= \alpha_{1/2} (\text{Tr } \Delta_R^\dagger \Delta_L + R \rightarrow L) \text{Tr } \bar{\Phi}^\dagger \Phi \\
&+ \alpha_{2/2} (\text{Tr } \Delta_R^\dagger \Delta_R + R \rightarrow L) (\text{Tr } \tilde{\Phi}^\dagger \Phi + \text{h.c.}) \\
&+ \alpha_{3/2} (\text{Tr } \Delta_R^\dagger \bar{\Phi}^\dagger \Phi \Delta_R + R \rightarrow L)
\end{aligned}$$

$$\begin{aligned}
\bar{\Phi} &\rightarrow U_L \bar{\Phi} U_R^\dagger, \quad \tilde{\Phi} \rightarrow U_L \tilde{\Phi} U_R^\dagger \\
\Phi^\dagger &\rightarrow U_R \Phi^\dagger U_L^\dagger \quad (\tilde{\Phi} \equiv \sigma_2 \bar{\Phi}^* \sigma_2)
\end{aligned}$$

What about $\text{Tr } \Delta_R^\dagger \tilde{\Phi}^\dagger \bar{\Phi} \Delta_R + \text{h.c.}$

+ R → L

$$\tilde{\Phi}^\dagger \bar{\Phi} = \left(\frac{1}{2} \text{Tr } \tilde{\Phi}^\dagger \bar{\Phi} \right) \mathbb{1}$$



$$T_\nu \Delta_R^\dagger \tilde{\Phi}^\dagger \Phi \Delta_R = \frac{1}{2} \underbrace{T_\nu \Delta_R^\dagger \Delta_R T_\nu \tilde{\Phi}^\dagger \Phi}_{\alpha_2 \text{ term}}$$



$V_{\Phi \Delta} = \text{complete!}$

\downarrow mass spectrum of Φ

mass terms + quanta $\Delta_R \rightarrow \langle \Delta_R \rangle$



$$T_\nu \tilde{\Phi}^\dagger \tilde{\Phi} = T_\nu \Phi^\dagger \Phi$$

$$\begin{aligned}
 V_{\text{mass}}(\bar{\Phi}) &= -\frac{\mu_0^2}{2} T_\nu \bar{\Phi}^+ \bar{\Phi} - \frac{\mu_0^2}{2} T_\nu \tilde{\Phi}^+ \bar{\Phi} \\
 &\quad + \frac{\alpha_1}{2} v_R^2 T_\nu \bar{\Phi}^+ \bar{\Phi} + \frac{\alpha_2}{2} v_R^2 T_\nu \tilde{\Phi}^+ \bar{\Phi} \quad (+ \text{h.c.})
 \end{aligned}$$

$$+ \frac{\alpha_3}{2} T_\nu \langle \Delta_R \rangle^+ \bar{\Phi}^+ \bar{\Phi} \langle \Delta_R \rangle$$

$$\langle \Delta_R \rangle = \begin{pmatrix} 0 & 0 \\ v_R & 0 \end{pmatrix}$$

$$\bar{\Phi} = \begin{pmatrix} \tilde{\phi}_1 & \phi_2 \end{pmatrix} = \begin{pmatrix} \phi_1^{0^+} & \phi_2^+ \\ -\phi_1^- & \phi_2^0 \end{pmatrix}$$

\Downarrow

$$T_\nu \langle \Delta_R \rangle^\dagger \bar{\Phi}^\dagger \Phi \langle \Delta_R \rangle =$$

$$= T_\nu \begin{pmatrix} 0 & \nu_R \\ 0 & 0 \end{pmatrix} (\bar{\Phi}^\dagger \Phi) \begin{pmatrix} 0 & 0 \\ \nu_R & 0 \end{pmatrix}$$

$$= (\bar{\Phi}^\dagger \Phi)_{22}$$

$$= \begin{pmatrix} \varphi_1^0 & -\varphi_1^+ \\ \varphi_2^- & \varphi_2^{0*} \end{pmatrix} \begin{pmatrix} \varphi_1^{0*} & \varphi_2^+ \\ -\varphi_1^- & \varphi_2^0 \end{pmatrix}$$

$$= (\varphi_2^- \varphi_2^+ + |\varphi_2^0|^2) = \underbrace{\varphi_2^+ \varphi_2}_{\text{SU}(2)_L \text{ DW.}}$$

$$\phi_i = \begin{pmatrix} \varphi_i^+ \\ \varphi_i^0 \end{pmatrix}$$

$$\mu^2 = \mu_0^2 + \alpha_1 V_R^2$$

$$\bar{\mu}^2 = \bar{\mu}_0^2 + \alpha_2 V_R^2$$

$$T_\nu \bar{\Phi}^+ \Phi = \phi_1^+ \phi_1 + \phi_2^+ \phi_2$$

(same masses)

$$T_\nu \tilde{\Phi}^+ \Phi = \phi_1^+ \phi_2 (\phi_2^+ \phi_1)$$

↑↑

$SU(2)_L$ point of view:

$$\phi_1^+ \phi_1, \quad \phi_2^+ \phi_2, \quad \phi_1^+ \phi_2 + \phi_2^+ \phi_1$$



$$V_{\text{mass}}(\Phi) = -\frac{\mu^2}{2}(\phi_1^\dagger \phi_1 + \phi_2^\dagger \phi_2)$$

$$- \frac{\bar{\mu}^2}{2}(\phi_1^\dagger \phi_2 + \phi_2^\dagger \phi_1)$$

$$+ \frac{\alpha_3}{2} v_R^2 \phi_2^\dagger \phi_2$$

$$\langle \Phi \rangle = \begin{pmatrix} v_1 & 0 \\ 0 & v_2 e^{i\alpha} \end{pmatrix}$$

$$(i) \quad v_2 = 0, \quad v_1 \equiv v$$



$$\bar{\Phi} = \begin{pmatrix} v+h+iG^0 & \phi_2^+ \\ -G^- & \phi_2^0 \end{pmatrix}$$

} mass analysis

$$\left(\mu^2 \simeq \bar{\mu}^2 \simeq M_W^2 (M_L^2) \right)$$

\Downarrow

$$M_h^2 \simeq M_W^2, \quad (G^-, G = \text{eaten})$$

$$M_{\phi_2}^2 \simeq \alpha_3 V_A^2$$

$\phi_2 = \text{heavy scalar doublet}$



only SM Higgs = "light"

Q. Was this automatic?

$\phi_1 = \text{light}$, $\phi_2 = \text{heavy}$
($\sim M_L$) ($\sim M_R$)

A. NO

$$\mu^2 = \mu_0^2 + \underbrace{\alpha_1 v_R^2}_{\text{large } (\sim v_R)} \approx M_L^2$$

$M_L \ll M_R \rightarrow$ limit: $M_L = 0$

illustration

$$\mu^2 = 0 \Rightarrow$$

$$\mu_0^2 + \alpha_1 v_R^2 = 0$$

$$\Rightarrow \mu_0^2 = -\alpha_1 v_R^2 \text{ (huge)}$$

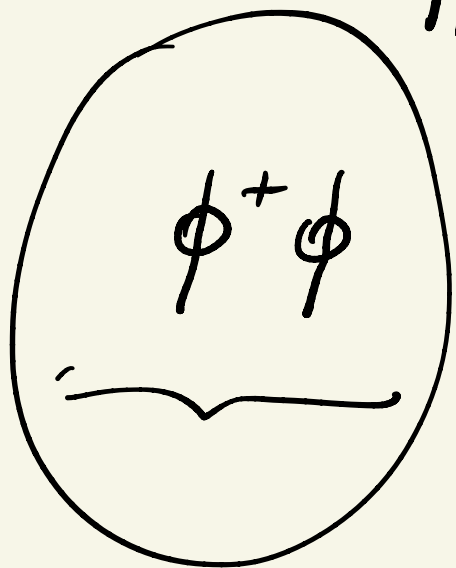
Fine-Tuning (FT)





Hierarchy Problem

$$M_L \ll M_R (?)$$



$$T, \Delta^+ \Delta$$

gauge invariant

limit: $m_\phi \rightarrow 0 \Rightarrow$

no new symmetry



$\phi^\dagger \phi =$ invariant under

$$\phi \rightarrow U \phi$$



scalar mass \neq protected
by a symmetry



completely different from
fermion mass

Fermion mass

$$m_f \bar{f} f = m_f f^\dagger \gamma_0 f$$



gauge charges

$$\bar{f} f = f^\dagger \gamma_0 f = \bar{f}_L f_R + \bar{f}_R f_L$$

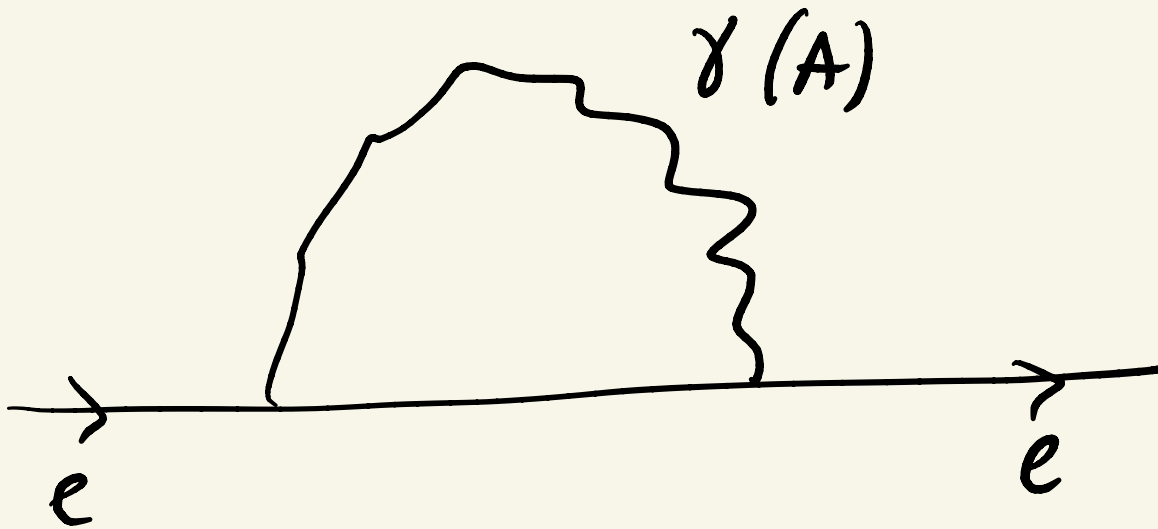
• $m_f = 0$ limit

$$\Rightarrow \mathcal{L}_D = i \bar{f}_L \gamma^\mu \partial_\mu f_L + i \bar{f}_R \gamma^\mu \partial_\mu f_R$$

fermion mass = protected
by chiral symmetry

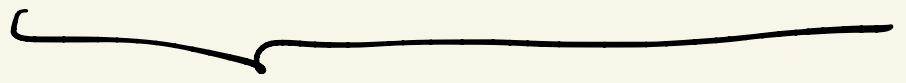
Veislopf 1930

loop correction to e mass



\Downarrow $W_e^0 (\bar{e}_L e_R + h.c.)$

$$\mu_e = \mu_e^0 \left[1 + \frac{\alpha}{\pi} \ln \frac{\Lambda}{\mu_e^0} \right]$$



L loop

145 th loop

↯ claim

$$\mu_e = \mu_e^0 \left[1 + \left(\frac{\alpha}{\pi} \right)^{145} \ln \frac{\Lambda}{\mu_e^0} \right]$$

1-loop

$$\mu_e = \mu_e^0 \left[1 + \frac{\alpha}{\pi} \ln \frac{\Lambda}{\mu_e^0} \right]$$

and not

$$M_e = m_e^0 + \frac{\alpha}{\pi} \Lambda^2$$

cannot
be

as in the scalar

once again: $m_e^0 \rightarrow 0$



chiral symmetry: $f_A \rightarrow e^{i\gamma_5} f_A$
 $f_L \rightarrow f_L$



pert. theory (to all orders)

must have : $f_a \rightarrow e^{i\vec{r}} f_a, t_L \rightarrow t_L$



$$w_e \text{ (to all orders)} = 0$$

$$\Rightarrow w_e \propto w_e^0 \dots$$



$$w_e = w_e^0 \left[1 + \left(\frac{d}{\pi} + \left(\frac{d}{\pi} \right)^2 \dots \right) \ln \frac{1}{w_e^0} \right]$$

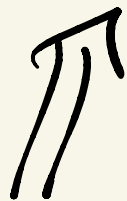
divergent symmetry =

= protective symmetry

Scalars

$$\mu^2 = \mu_0^2 + \frac{\alpha}{\Lambda} \Lambda^2 + \dots$$

$$(\mu^2 = \mu_0^2 + \alpha V_R^2)$$



LRSM

FT = ugly, but
not a problem

• log. "infinity" = small

$$\Lambda \leq M_{\text{pl}} = G_N^{-1/2}$$

↑
string gravity

$$\left(\frac{\alpha}{4\pi} \ln \frac{M_{\text{pl}}}{m_e} \ll 1 \right)$$

"naturalness" (technical sense)



't Hooft '70s

$\emptyset \rightarrow 0 \Rightarrow$ new symmetry



¹
a parameter

$$\Theta^{\text{loop}} = \Theta_0 [1 + \dots]$$

- fermion mass = "naturally" small
(protected)

- scalar mass \neq naturally small

\Leftrightarrow not protected

Q. Go back to \mathbb{F}

$$v_2 = 0 \Leftrightarrow \phi_2 \neq \text{Higgs}$$

$$\Downarrow$$
$$\mu_{\phi_2}^2 \simeq v_R^2$$

but:

$\alpha_3 v_R^2 \phi_2^\dagger \phi_2$ and ϕ_2 is

in general a part of Higgs
($v_2 \neq 0$)

$$\downarrow v_2 \neq 0$$

$$\phi_1, \phi_2 \rightarrow (\langle \phi_i \rangle = v_i)$$

$$h = \cos\beta \phi_1 + \sin\beta \phi_2$$

$$H = -\sin\beta \phi_1 + \cos\beta \phi_2$$

$$\tan\beta \equiv v_2/v_1 \quad \dots$$

$$\langle h \rangle = \sqrt{v_1^2 + v_2^2} \quad (\text{Higgs})$$

$$\langle H \rangle = 0 \quad (\text{scalar doublet})$$

$$\Rightarrow \boxed{M_H^2 \approx \alpha_3 v_R^2 \frac{1}{\cos 2\beta}}$$

$m_h^2 \stackrel{!}{=} \text{small by FT}$

Fine Tuning = necessary