


BB SM Neutrino Course

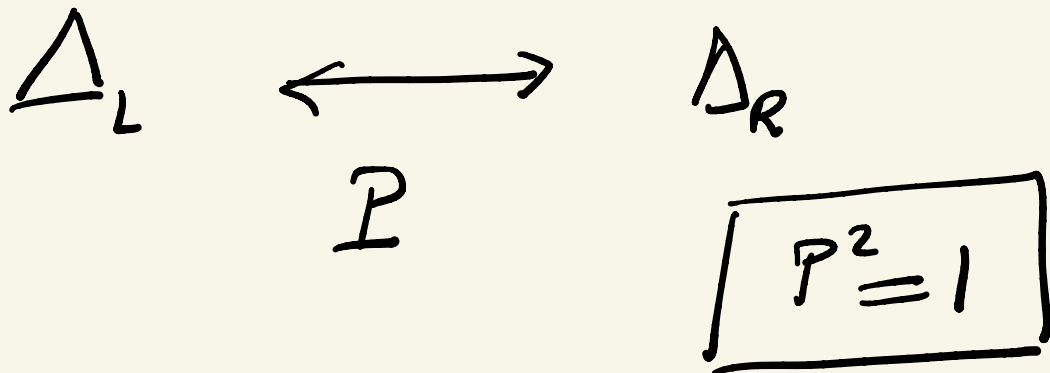
Lecture XXII

L.M.-U

Spring 2020



Domain Walls



$$V = -\frac{\mu^2}{2} (\Delta_L^2 + \Delta_R^2) + \lambda (\Delta_L^2 + \Delta_R^2)^2$$

$+ \lambda' \Delta_L^2 \Delta_R^2$

$\lambda' > 0$ $\Rightarrow \langle \Delta_L \rangle = 0, \langle \Delta_R \rangle \neq 0$

$\Delta = \text{triplets under } SU(2)$

$\langle D_L \rangle = 0 \rightarrow$ concentrate on Δ_R

$$V = -\frac{\mu^2}{2} T_\nu \Delta_R^\dagger \Delta_R + \lambda_1 (T_\nu \Delta_R^\dagger \Delta_R)^2$$

$$+ \lambda_2 T_\nu \Delta^2 T_\nu \Delta^{\dagger 2}$$

$$T_\nu \Delta_R^\dagger \Delta_R \Delta_R^\dagger \Delta_R = f(\dots)$$

$T_\nu \Delta_R^\dagger \Delta_R =$ does not care who gets

\Downarrow $\nu e \nu$

$$\Delta_R = 3 \text{ complex} = 6 \text{ real}$$

\Downarrow

$SO(6)$ symmetry

$$\lambda_2 = 0: SO(6) \rightarrow SO(5)$$

$$\Rightarrow \frac{6 \cdot 5}{2} - \frac{5 \cdot 4}{2} = 5$$

broken generators

\Rightarrow 5 NG bosons

3 creators: w_R^+ , w_R^- , z_R

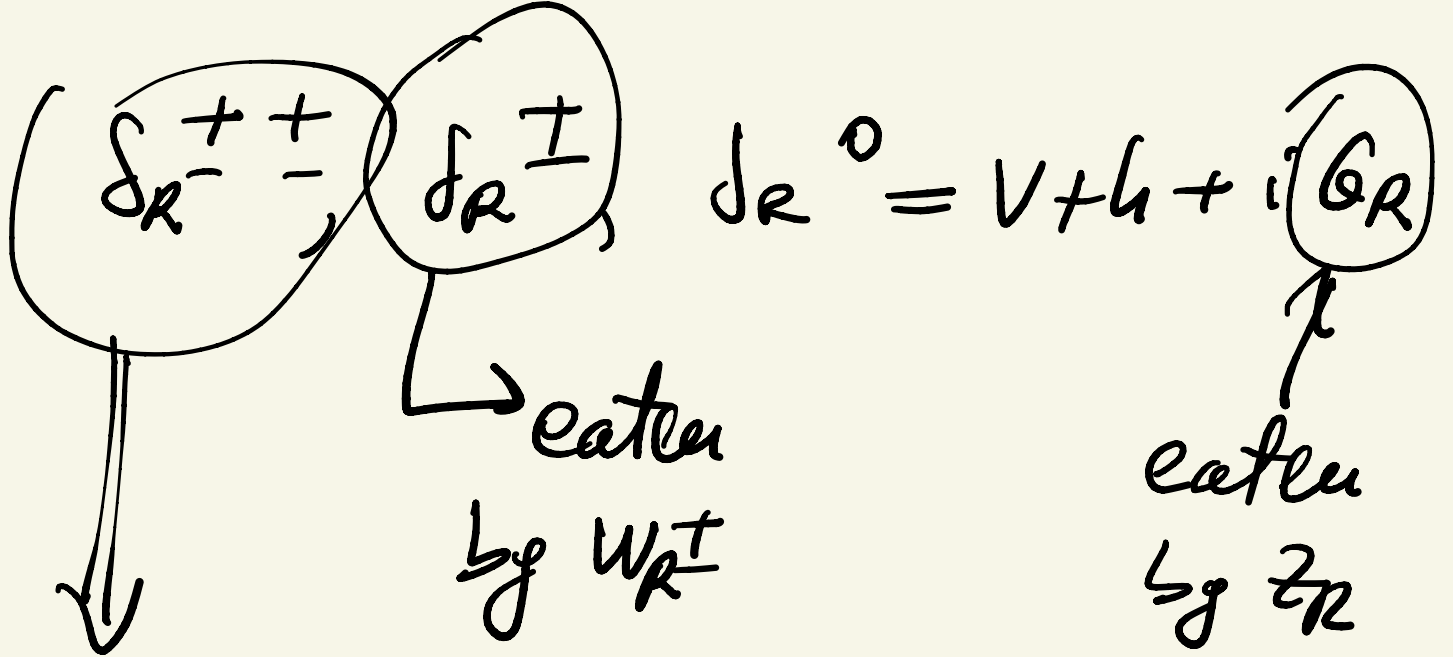
\Downarrow

2 NG

$$\lambda_2 \neq 0 \quad \boxed{\lambda_2 > 0}$$

$$\langle \Delta_R \rangle = \begin{pmatrix} 0 & 0 \\ v_R & 0 \end{pmatrix}$$

local minimum



$$M_\delta^{++} \propto \lambda_2 \phi_R^2$$

$\lambda_2 = 0 \Rightarrow 2 \text{ NG bosons}$

$$SU(2)_R \times U(1)_{B-L}$$

gauge \downarrow $4 - 1 = 3$ broken

$$U(1)_Y$$



3 massive W_R^\pm, Z_R

go back to SM

$$SU(2) \times U(1) \text{ gauge}$$

$$V \Rightarrow SO(4) \text{ global}$$

$$\begin{array}{c} | \\ \underbrace{SU(2)_L \times SU(2)_R}_{\text{gauge}} \quad \underbrace{\hspace{2cm}}_{\text{part gauge}} \end{array}$$

• gauge: $SU(2) \times U(1) \rightarrow U(1)$

$$3 + 1 - 1 = 3$$

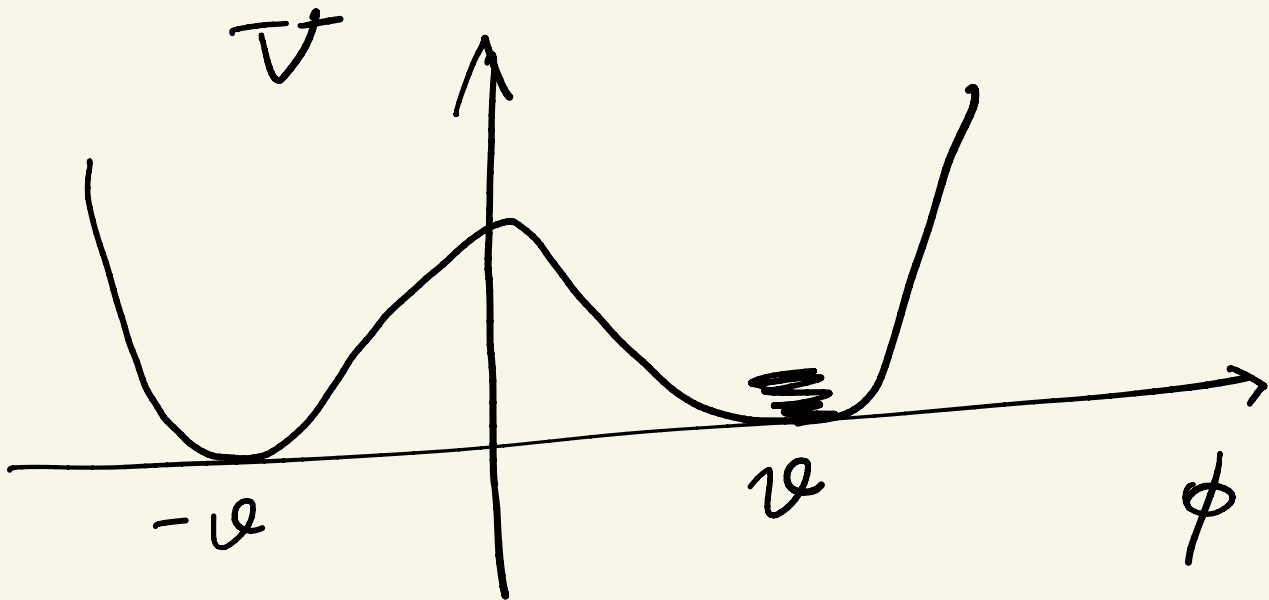
• global $SO(4) \rightarrow SO(3)$

$$6 - 3 = 3$$

$$D = \mathbb{Z}_2 \text{ discrete}$$

$$\mathcal{L} = \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - V(\phi)$$

$$\phi \in \mathbb{R} \quad V = \frac{1}{2} (\phi^2 - v^2)^2$$

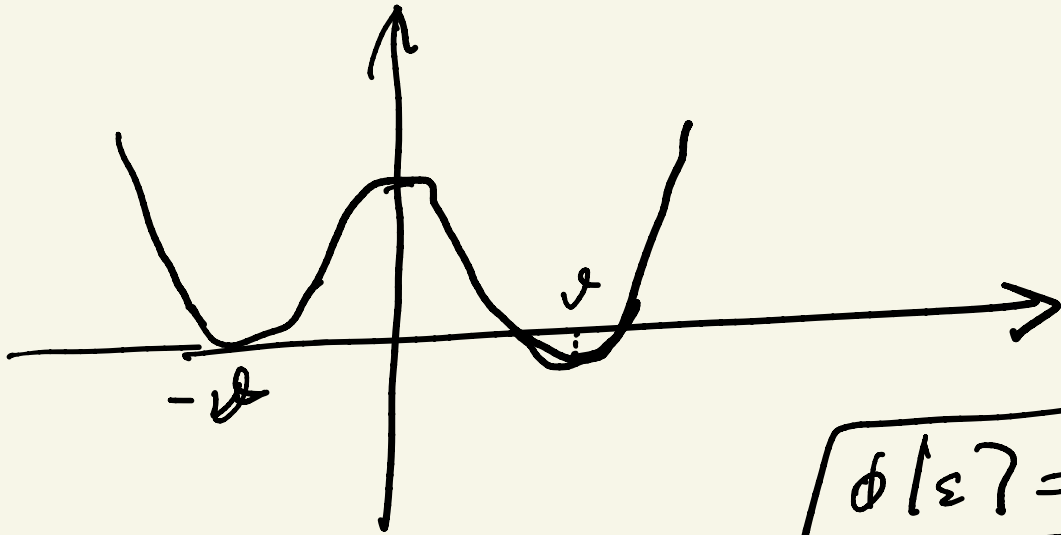


No sym. breaking in QM

\Rightarrow tunneling

$$\text{QM: } \Gamma \propto e^{-\int_a^b dx \sqrt{V}}$$

Coleman 74



$$\phi[\varepsilon] = (u)^\dagger$$

$$V = (\phi^2 - v^2)^2 + \frac{\varepsilon}{v} (\phi - v)$$

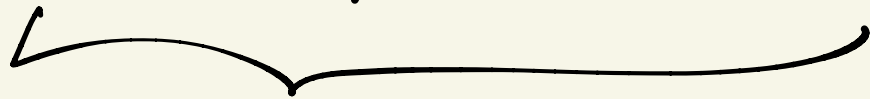
$$\Gamma_{\text{turn}} \propto e^{-v^{12}/\varepsilon^3}$$

$$\begin{aligned} M_0 &= \{ \phi_0 : V = \min \} = \{ \phi_0^2 = v^2 \} \\ &= \{ \pm v \} = \mathbb{Z}_2 \\ &\quad \text{vacuum manifold} \end{aligned}$$

$$\phi_0^E = \psi, \quad \phi_0^H = -\psi$$

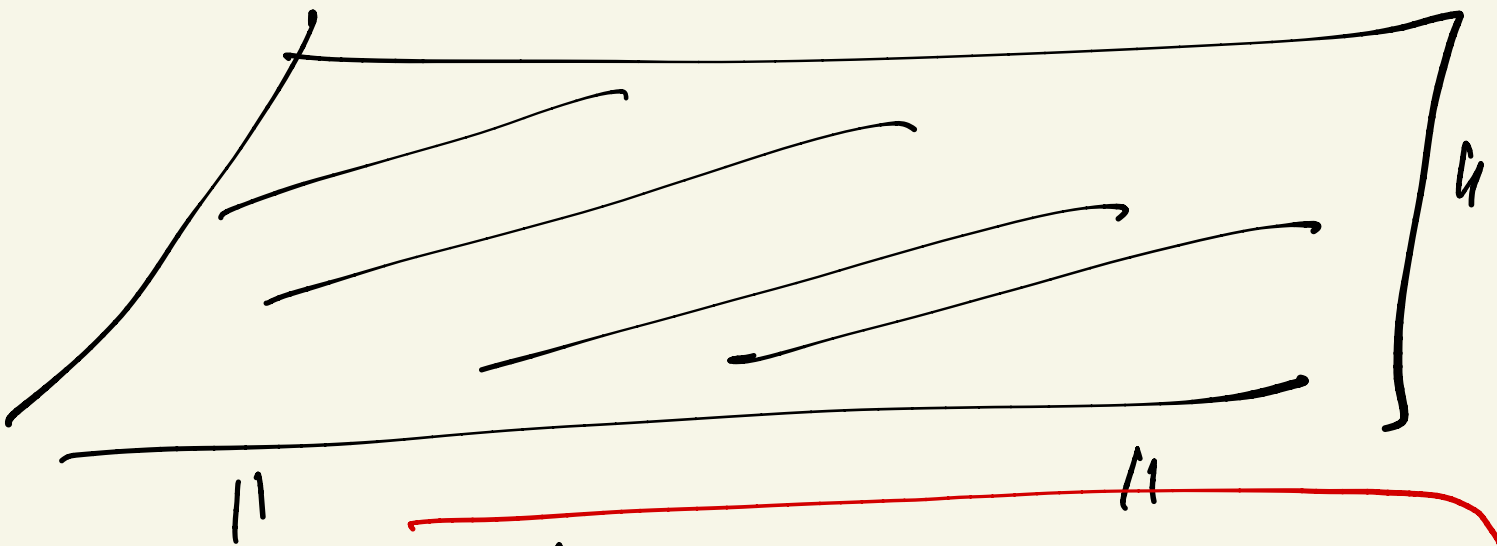


$$\bar{V}_0 = 0, \quad K_E = 0$$



$E=0$ ground state

$\phi_{cl} = \phi_{dw} = \phi_w =$ arbitrary f-u
of z



$S \rightarrow \infty$ well (x-y)
 $\phi_w(z)$ static solution

Finite energy (density)

$$\underbrace{E/S}_{\text{finite}} = \int_{-\infty}^{+\infty} dz \left[\underbrace{\frac{1}{2} \left(\frac{d\phi}{dz} \right)^2}_0 + \underbrace{V}_{0} \right]$$

at ∞

$$\mathcal{M}_\infty = \{ z = +\infty, -\infty \}$$

$$\mathcal{M}_0 = \{ +v, -v \}$$

$$\mathcal{M}_\infty : \left. \begin{array}{l} V \rightarrow 0 \\ \frac{d\phi}{dz} \rightarrow 0 \end{array} \right\} \Rightarrow \boxed{\phi_\infty \Rightarrow \pm v}$$

map from \mathcal{M}_∞ to \mathcal{M}_0

(1) trivial map

$$\phi_0(z=+\infty) = \phi_0(z=-\infty) = +v$$

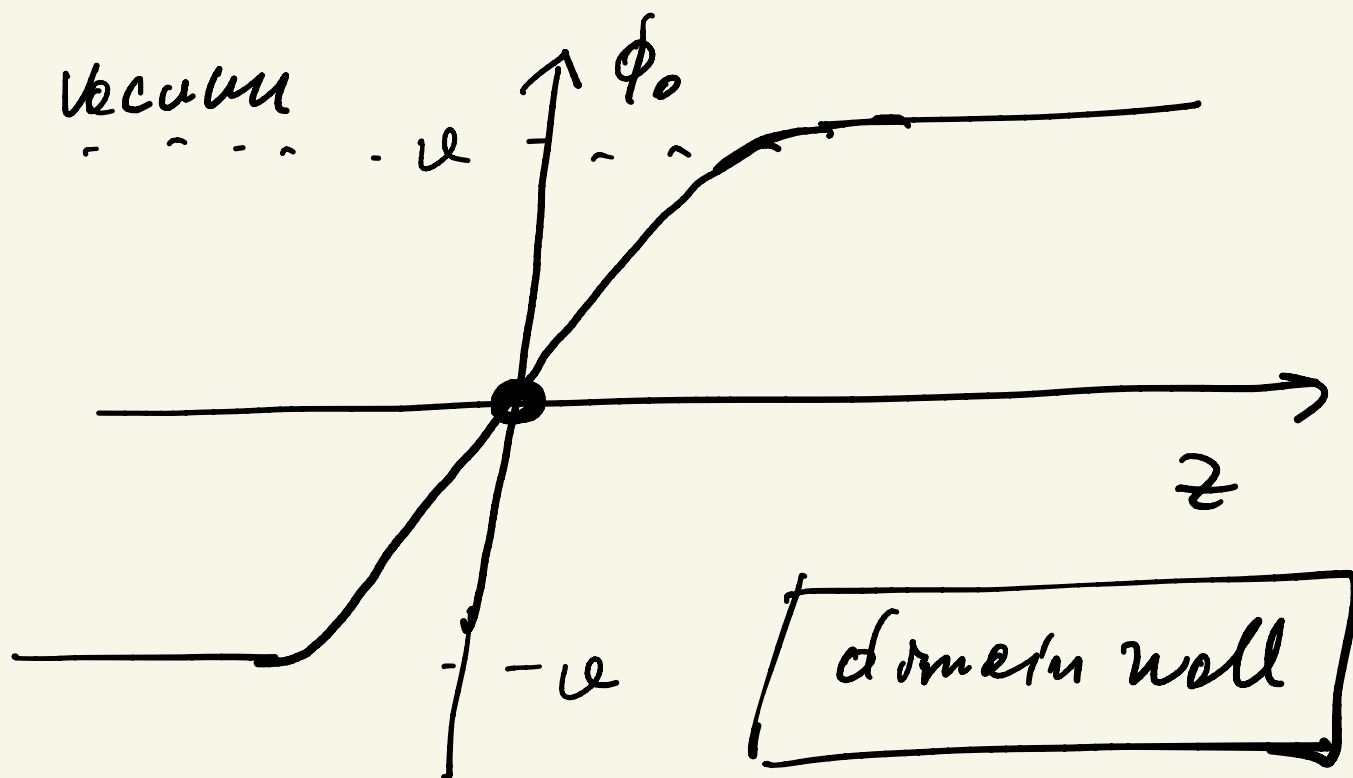
$$\mathcal{M}_\infty \rightarrow +v$$

solution = vacuum

(ii) non-trivial case

$$(a) \quad \phi_0(z = +\infty) = +v$$

$$\phi_0(z = -\infty) = -v$$



(b)

$$\phi_0(z = +\infty) = -v$$

$$\phi_0(z = -\infty) = +v$$

anti-wall

$$\mathcal{L} = \frac{1}{2}(\partial_\mu \phi)^2 - V(\phi)$$

$$\square \phi = - \frac{\partial V}{\partial \phi} \quad (1)$$

Static

$$\frac{d^2 \phi_w}{dz^2} = \frac{\partial V}{\partial \phi} \cdot \frac{d\phi_w}{dz}$$

$$\frac{1}{2} \frac{d}{dz} \left(\frac{d\phi_w}{dz} \right)^2 = \frac{dV}{dz} \quad (2)$$

$$\frac{1}{2} \left(\frac{d\phi_w}{dz} \right)^2 = V + \text{const.} \quad (3)$$

at ∞ : \downarrow \downarrow finite energy

$$\Rightarrow \text{const.} = 0$$



$$\frac{d\phi_w}{dz} = \pm \sqrt{2V} \quad (4)$$

$$V = \frac{1}{2} (\phi_w^2 - a^2)^2$$



$$\frac{d\phi_w}{dz} = \pm (\phi_w^2 - a^2)$$



$$\frac{d\phi_w}{\phi_w^2 - a^2} = \pm dz \quad (5)$$



$$\left. \begin{aligned} \phi_w &= v \tanh vz \\ \phi_{out w} &= -v \tanh vz \end{aligned} \right\} (6)$$

Thickness $\delta_w = \frac{1}{v}$

$$\begin{aligned} E/s &= \int_{-v}^{+v} dz \left[\frac{1}{2} \left(\frac{d\phi_w}{dz} \right)^2 + \overline{V} \right] \\ &= \int_{-v}^{+v} d\phi \frac{dz}{d\phi} 2\overline{V} = \int_{-v}^{+v} d\phi \sqrt{2\overline{V}} \end{aligned}$$

$$\left(\frac{d\phi}{dz} = \sqrt{2\overline{V}} \right)$$

$$\frac{E}{S} = \pm \int_{-v}^{+v} d\phi (\phi^2 - v^2)$$

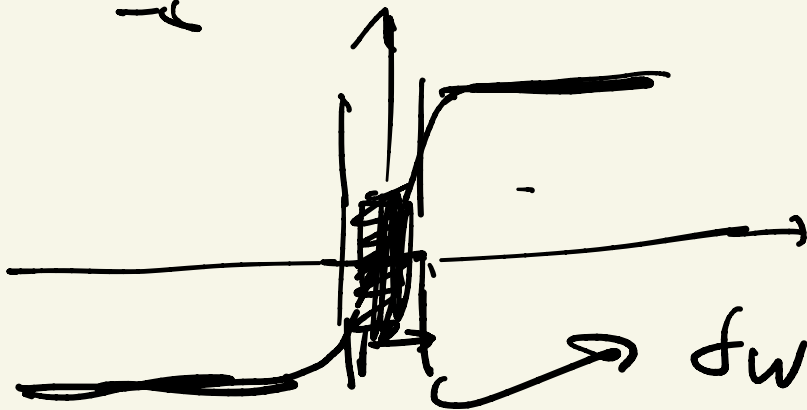
$$\approx \pm v^3$$

$$E \approx v^3 S$$

Energetics

$$\frac{E}{S} = \int_{-z}^{+z} dz \left[\frac{1}{2} \left(\frac{d\phi}{dz} \right)^2 + V(\phi) \right]$$

*





$$V(\phi) = \frac{1}{2}(\phi^2 - v^4)^2 = \frac{1}{2}v^4 \text{ at } \phi=0$$

$V(\phi)$ piece $\frac{1}{2}v^4 \int_{-\delta}^{\delta} = v^4 \delta$

* $\text{minimum} \rightarrow \delta \rightarrow 0$

gradient $\frac{1}{2} \left(\frac{\Delta \phi_w}{\Delta z} \right)^2 \cdot 2\delta$

$$= \frac{1}{2} \left(\frac{2v}{2\delta} \right)^2 2\delta = \frac{v^2}{\delta}$$

$$E/\delta = v^4 \delta + v^2/\delta$$

$$\frac{\partial V}{\partial \delta} = 0 \Rightarrow v^4 = v^2/\delta^2$$

$$f = \frac{1}{2\epsilon}$$

$$j_\mu = \epsilon_{\mu\nu} \partial^\nu \phi \quad \boxed{\mu, \nu = t, z}$$

$$\epsilon_{\mu\nu} = -\epsilon_{\nu\mu}$$

$$\partial^\nu j_\mu = 0 = \epsilon_{\mu\nu} \partial^\nu \partial^\mu \phi$$

$$j_0 = \frac{d\phi}{dt} \Rightarrow$$

$$Q = \int_{-\infty}^{+\infty} dt \frac{d\phi}{dt}$$

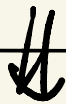
$$Q = \phi(+\infty) - \phi(-\infty)$$

$$\frac{dQ}{dt} = 0$$

$$Q(\text{vacuum}) = 0 ;$$

$$Q(\text{well}) = 2\vartheta ;$$

$$Q(\text{anti-well}) = -2\vartheta$$



well is stable!

• well + anti-well \rightarrow vacuum

+
domein well

universe

Kobzarev, Okun
Zeldovich '70

$$E/\rho = v^3$$

$$v \geq 1 \text{ TeV}$$

$$E_{\text{visible}} \approx \frac{1}{10} E_{\text{universe}}$$

$$E_{\text{univ}} \approx 10 E_{\text{visible}}$$

$$\frac{n_B \rightarrow \text{baryons}}{n_\gamma \rightarrow \text{radiation}} = 10^{-10}$$

$$E_\gamma \approx 0 \text{ K} \approx 10^{-4} \text{ eV}$$

$$E_B \approx 6 \text{ GeV} \approx 10^9 \text{ eV} \approx 10^{13} E_\gamma$$

$$E_{\text{new}} = 10 E_B = 10 N_B E_B$$

$$= 10 \cdot 10^{-10} N_\gamma \text{ GeV}$$

$$= 10 \cdot 10^{-10} (R^3 \cdot n_\gamma) \text{ GeV}$$

$$E_{\text{uni}} = 10^{-9} (RT)^3 \text{ GeV}$$

$$E_{\text{well}} = v^3 R^2$$

$$T = 10^{-4} \text{ eV}$$

$$R = 10^{28} \text{ cm}$$

$$T = 10^{-13} \text{ GeV}$$

$$RT = 10^{28} \cdot 10^{-13} \text{ cm GeV}$$

⊕
proton

$$\text{GeV}^{-1} \approx 10^{-14} \text{ cm} \Rightarrow \text{GeV cm} \approx 10^{14}$$

$$RT = 10^{29}$$

$$\Downarrow RT = \text{const.}$$

$$\frac{E_{\text{well}}}{E_{\text{unv}}} \approx \frac{v^3 R^2}{10^{-9} (RT)^3 \text{ GeV}} \approx$$

$$\approx \frac{10^9 \text{ GeV}^3 R^2}{10^{-9} \cdot 10^{87} \text{ GeV}}$$

$$\approx 10^{-69} (\text{GeV} R)^2$$

$$\approx 10^{-69} (\text{GeV} 10^{28} \text{ cm})^2$$

$$\approx 10^{-69} (10^{42})^2 \approx 10^{15} !!!$$

Standard cosmological model

U big bang

1. cold today
2. hot yesterday
3. isotropic and homogeneous

$$d_u = t$$

$$t \approx \frac{1}{T} \frac{M_{pl}}{T} \leftarrow \underline{\underline{gravity}}$$

⇓

fast

$$R(T) = \frac{c}{T}$$

$$c = 10^{29}$$

$$d_H(t) = \frac{H_{pe}}{T^2}$$

size

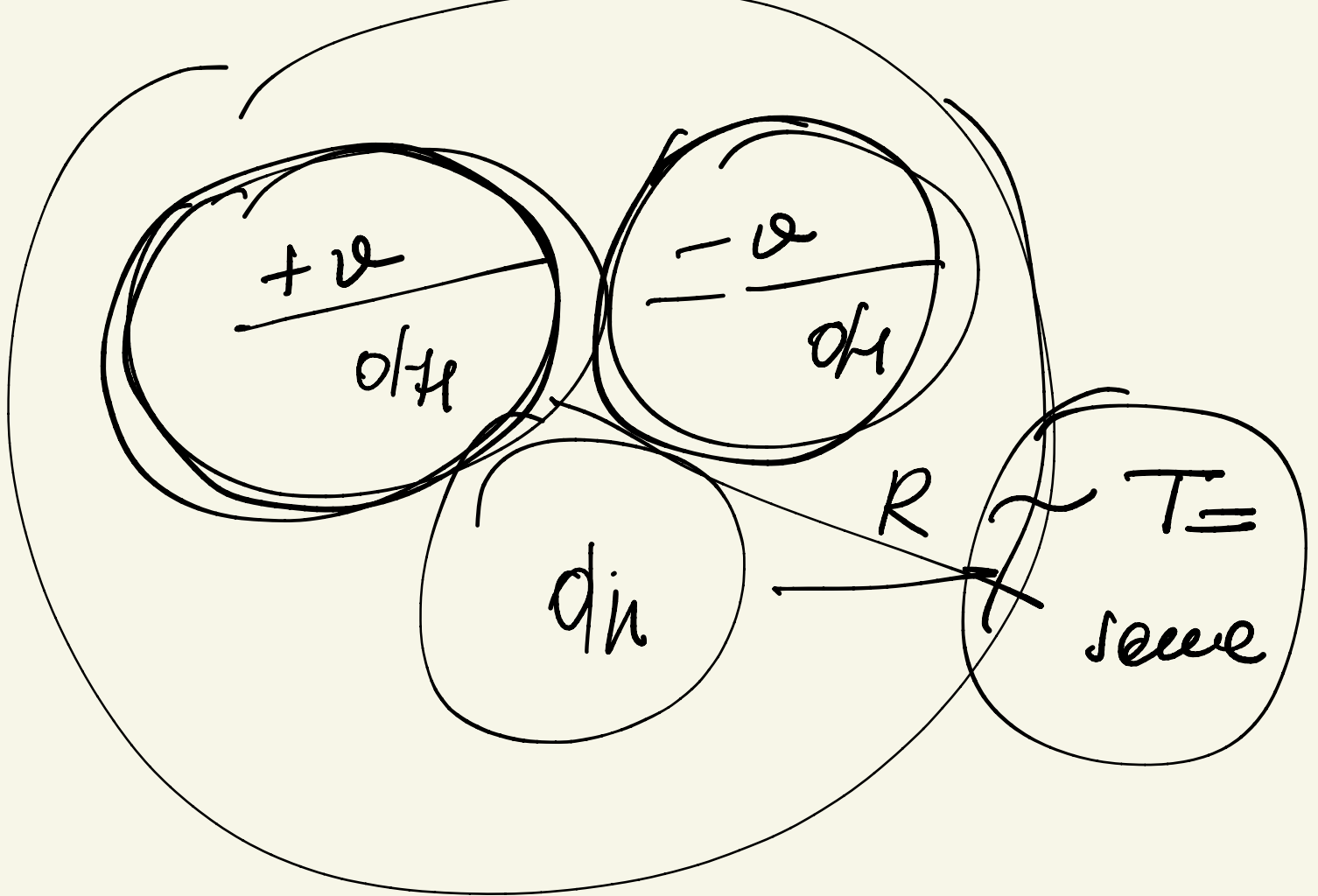


$$\frac{R(T)}{d_H(t)} = c \frac{T}{H_{pe}}$$

cancel contact

$$\frac{R_0}{d_{H^0}} = 1 \text{ (today)}$$

$$R(T) \geq d_H(t)$$



$$c = MeV = 10^{-3} \text{ GeV}$$

$$R/dh \approx 10^{29} \cdot \frac{10^{-3}}{10^{+18}} = 10^8$$